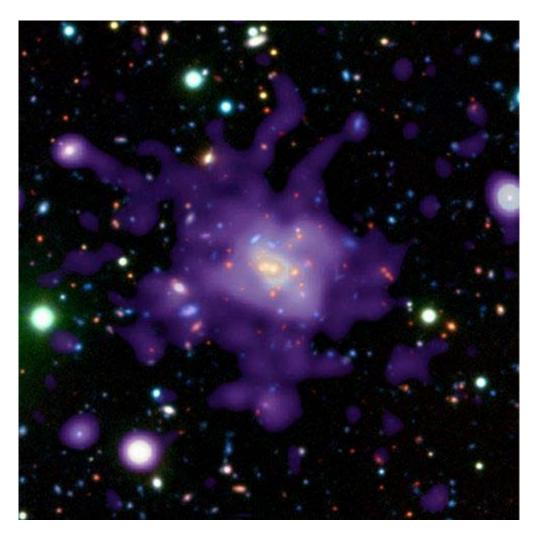
Movement of Galaxy Clusters Opens New Window on the Universe -- May Unlock Key to Dark Energy

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A large research team from two major astronomy surveys reports that scientists detected the movement of distant galaxy clusters via the kinematic <u>Sunyaev-Zel'dovich</u> (kSZ) effect, which has never before been seen.

First proposed in 1972 by Russian physicists <u>Rashid Sunyaev</u> and <u>Yakov</u> <u>Zel'dovich</u>, the kSZ effect results when the hot gas in galaxy clusters distorts the cosmic microwave background radiation—which is the glow of the heat left over from the Big Bang—that fills our universe. Radiation passing through a galaxy cluster moving toward Earth appears hotter by a few millionths of a degree, while radiation passing through a cluster moving away appears slightly cooler.

Pedro Ferreira, an astrophysics professor at the University of Oxford, calls the paper a "beautiful piece of work" that neatly demonstrates an accurate method for studying the evolution of the universe and the distribution of matter in it. Ferreira had no role in the research but is familiar with it. "This is the first time the kSZ effect has been unambiguously detected, which in and of itself is a really important result," Ferreira says. "By probing how galaxies and clusters of galaxies move around in the universe, the kSZ effect is directly probing how objects gather and evolve in the universe," he says. "Therefore it is hugely dependent on dark matter and dark energy. You can then think of the kSZ effect as a completely new window on the large-scale structure of the universe."

Now that it has been detected, the kSZ effect could prove to be an exceptional tool for measuring the velocity of objects in the distant universe, the researchers report. It could provide insight into the strength of the gravitational forces pulling on galaxy clusters and other bodies.

Chief among these forces are the still-hypothetical dark energy and dark matter, which are thought to drive the universe's expansion and the motions of galaxies. In addition, the strength of the kSZ effect's signal depends on the distribution of electrons in and around galaxies. As a result, the effect also can be used to trace the location of atoms in the nearby universe, which can reveal how galaxies form.

The benefits of the kSZ effect stem from a unique ability to pinpoint velocity, says Hand, a 2011 Princeton graduate who is now a graduate student in astronomy at the University of California-Berkeley. The researchers detected

the motion of galaxy clusters that are several billion light years away moving at velocities of up to 600 kilometers (372 miles) per second.

"Traditional methods of measuring velocities require very precise distance measurements, which is difficult. So, these methods are most useful when objects are closer to Earth," Hand says.

"One of the main advantages of the kSZ effect is that its magnitude is independent of a galaxy cluster's distance from us, so we can measure the velocity of an object's motion toward or away from Earth at much larger distances than we can now," Hand says. "In the future, it can provide an additional statistical check that is independent of our other methods of measuring cosmological parameters and understanding how the universe forms on a larger scale."

To find the kSZ effect, the researchers combined and analyzed data from the ACT and BOSS projects. The kSZ effect is so small that it is not visible from the interaction with an individual galaxy cluster with the cosmic microwave background (CMB), but can be detected by compiling signals from several clusters, the researchers discovered.

ACT is a custom-designed 6-meter telescope in Chile built to produce a detailed map of the CMB using microwave frequencies. The ACT collaboration involves a dozen universities with leading contributions from Princeton and the University of Pennsylvania, and includes important detector technology from NASA's Goddard Space Flight Center, the National Institute of Standards and Technology, and the University of British Columbia.

BOSS, a visible-light survey based at the <u>Apache Point Observatory</u> in New Mexico, has captured spectra of thousands of luminous galaxies and quasars to improve understanding of the large-scale structure of the universe. BOSS is a part of the Sloan Digital Sky Survey III, the third phase of the most

productive astronomy project in history, and a joint effort among 27 universities and institutions from around the world.

For the current project, researchers from ACT compiled a catalog of 27,291 luminous galaxies from BOSS that appeared in the same region of sky mapped by ACT between 2008 and 2010. Because each galaxy likely resides in a galaxy cluster, their positions were used to determine the locations of clusters that would distort the <u>CMB radiation</u> that was detected by ACT.

Hand used the 7,500 brightest galaxies from the BOSS data to uncover the predicted kSZ signal produced as galaxy clusters interacted with CMB radiation. ACT collaborator Arthur Kosowsky, an associate professor of physics and astronomy at the University of Pittsburgh, suggested a particular mathematical average that reflects the slight tendency for pairs of galaxy clusters to move toward each other due to their mutual gravitational attraction, which made the kSZ effect more apparent in the data.

The overlap of data from the two projects was essential because the amplitude of the signal from the kSZ effect is so small, says ACT collaborator <u>David Spergel</u>, professor and department chair of astrophysical sciences at Princeton, as well as Hand's senior thesis adviser. By averaging the ACT's CMB maps with thousands of BOSS galaxy locations, the kSZ signal got stronger in comparison to unrelated signals and measurement errors, Spergel says.

"The kSZ signal is small because the odds of a microwave hitting an electron while passing through a galaxy cluster are low, and the change in the microwave's energy from this collision is slight," says Spergel. "Including several thousand galaxies in the dataset reduced distortion and we were left with a strong signal."

In fact, if analyzed separately, neither the ACT nor the BOSS data would have revealed the kSZ effect, Kosowsky says. "This result is a great example of an important scientific discovery relying on the rich data from more than one large astronomy survey," he says. "The researchers of the ACT and BOSS

collaborations did not have this in mind when they first designed their experiments."

The paper was initiated at <u>Princeton University</u> by lead author Nick Hand. Fifty-eight collaborators from the <u>Atacama Cosmology Telescope</u> (ACT) and the <u>Baryon Oscillation Spectroscopic Survey</u> (BOSS) projects are listed as coauthors. The paper was submitted to the journal Physical Review Letters. The image at the top of the page is a color composite image of the galaxy cluster RDCS 1252.9-2927 shows the X-ray (purple) light from 70-million-degree Celsius gas in the cluster, and the optical (red, yellow and green) light from the galaxies in the cluster. X-ray data from Chandra and the XMM-Newton Observatory show that this cluster was fully formed more than 8 billion years ago, and has a mass at least 200 trillion times that of the Sun.

At a distance of 8.5 billion light years, it is the most massive cluster ever observed at such an early stage in the evolution of the universe. Even though the cluster is seen as it was only 5 billion years after the Big Bang, it has an abundance of elements such as silicon, sulfur, and iron similar to that of clusters observed at more recent epochs. The cluster gas must have been enriched by heavy elements synthesized in stars and ultimately ejected from the galaxies. The relative abundances of these heavy elements are indicators of the star formation history of the galaxies. The observations of RDCS 1252.9-2927 are consistent with the theory that most of the heavy elements were produced by massive stars some 11 billion years ago.